

1 **The structure of the Azores triple junction: implications for S. Miguel Island**

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7 **Abstract**

8 The lack of reliable morphological and geophysical data for most of the Azores plateau
9 has been up to now a major limitation to the understanding of tectonic and magmatic
10 processes that shape the Eurasia-Nubia-North America triple junction and the Terceira
11 Rift. This situation changed recently: for the first time both the Triple Junction area and
12 the Terceira Rift are covered by high quality swath bathymetry surveys and marine
13 magnetic data with GPS quality positioning. This provides a good description of the
14 surface morphology, and also of magnetic chrons that give fundamental information for
15 the timing of spreading processes on the geological time frame. There is also a large
16 amount of data from continuous GPS stations, operating since 1997, which provides
17 accurate estimations of present day velocities for most of the islands. It is shown that
18 only two main rift systems can be found on the plateau, the oldest one matching the
19 Princess Alice Basin, and the newest one matching Terceira rift; the shift between the
20 two probably occurring close to ~3Ma. It is shown that extension is nowadays
21 concentrated in Terceira rift, progressively attaching Graciosa and Terceira islands to
22 Eurasia, while S. Miguel is being strained by rifting. It is also shown that no right lateral
23 strike slip fault connects Terceira Rift to the Mid-Atlantic Ridge and that the differential
24 motion between Eurasia and Nubia, west of Western Graciosa Basin, is accommodated
25 within a wide right lateral transtension area.

26 The succession of basins and volcanic highs that runs from the Western Graciosa Basin, close to
27 the Mid-Atlantic Ridge (MAR) up to Formigas islets, in the western tip of Gloria Fault (see
28 figure 1 for locations), was interpreted by Machado (1959) as a rift belt of tectonic origin and
29 referred to as the Terceira Rift. Krause and Watkins (1970) were the first to provide a plate
30 tectonic interpretation of the genesis of the Azores, describing it as a secondary spreading ridge,
31 based on the interpretation of coarse bathymetric and magnetic data. Its development was
32 attributed to a sudden change in the direction and the magnitude of the Eurasia-North America
33 and Nubia-North-America spreading rates, which would have occurred around 45 Ma BP.

34 McKenzie (1972) discussed the stability of such configuration, showing that it implies a
35 northward migration of the Eurasia-North America-Nubia triple point. Searle (1980), based on
36 additional shipborne magnetic data and a large amount of GLORIA side-scan sonar data
37 acquired between the Azores plateau and the MAR, was able to define the Terceira Rift more
38 fully. He proposed that the configuration change predicted by Krause and Watkins (1970)
39 corresponded to a northward jump of the EU-NU plate boundary from the latitude of the East
40 Azores Fracture Zones, to the vicinity of the North Azores Fracture Zone, which would connect
41 Terceira Rift to the MAR (Searle, 1980).

42 Srivastava *et al.* (1990) analysed a large magnetic compilation in the North Atlantic and
43 identified the main changes since chron M0 (~125 Ma, Cande & Kent, 1995). They showed that
44 Iberia was attached to Africa from the late Cretaceous to middle Eocene. Between the middle
45 Eocene to late Oligocene it behaved almost as an independent plate with slight motion relative
46 to Africa, while most of the deformation was concentrated along the northern border. Since the
47 late Oligocene Iberia is considered as a part of the Eurasia plate (Srivastava *et al.*, 1990). Luis
48 and Miranda (2008) refined the magnetic compilation of the North Atlantic, and concluded that
49 the establishment of the EU-AF-NA triple junction south of the Azores occurred between
50 chrons C11- C12 (ca. 30 Ma) and C6c (ca. 24 Ma) and following the addition of Iberia to
51 Eurasia.

52 The location and northward motion of the Azores triple point was analyzed by Luis *et al.* (1994)
53 at a segment scale. The study was based on a detailed aeromagnetic data set covering mostly the
54 area close to the MAR, up to chron c5 (~10 Ma). They concluded that the Eurasia-Nubia-North
55 America triple junction was located north of the East Azores Fracture Zone between anomaly 4
56 and anomaly c3a times, approximately at 38°20'N, 30°15'W (Eurasia fixed co-ordinates) and
57 proposed a present-day location close to 38°55'N, 30°00'W (Eurasia fixed co-ordinates), after
58 anomaly Chron c2a, approximately 2.45 Ma ago. Being so, the evolution of the Azores plateau
59 was determined first by the joining of Iberia to Eurasia and then by the progressive attachment
60 of Eurasia to Nubia (Lourenço *et al.*, 1998), associated with the northward migration of the
61 triple junction, and rifting and eventually the spreading associated with Terceira rift. The
62 reanalysis of regional marine magnetic data for the North-Atlantic also showed that the width of
63 the lithospheric sliver, that can be attributed to Azores rifting and incipient spreading, is much
64 smaller than the plateau itself, corresponding to a total extension close to 75 km. Stretching of
65 the Azores plateau took place essentially in the last ~20-25 Ma (Luis and Miranda, 2008).

66 The precise location and the nature of the geological processes associated with Azores extension
67 are more difficult to establish: the identification of young magnetic anomalies with a trend
68 similar to the Terceira rift was made by Miranda *et al.* (1991) and Luis *et al.* (1994); however,
69 no stable magnetic striping was found and apparently only Brunhes and Matuyama chrons can

70 be identified, and in some cases not strictly confined to the Terceira Rift: S. Jorge or the Pico-
71 Faial Islands also show large positive magnetic anomalies and there is evidence for the
72 Brunhes-Matuyama transition in the island of Faial, with clear continuation offshore. The fact
73 that only young chrons are represented in the new Azores domain was checked independently in
74 the Islands of Faial (Miranda *et al.*, 1991), S. Miguel (Johnson *et al.*, 1998) and S. Jorge (Silva
75 *et al.*, 2012) and this shows that spreading related with the Azores axis is young, ~2 Ma, and
76 driven by discontinuous stages of volcanism and tectonic movements, both in space and in time.
77 Gravity modeling (Luis *et al.*, 1998, Luis and Neves, 2006) confirmed that the secondary
78 Azores spreading ridge is not sufficiently mature to generate a significant thermal effect. Vogt
79 and Jung (2003) compared the Terceira Rift with other ultra-slow spreading ridges (e.g.
80 Southwest Indian and Gakkel) suggesting the existence of a hyper-slow regime characterized
81 by a wider rift valley and a thicker axial lithosphere, that would characterize the Azores.

82 The existence of a hotspot as a driving mechanism for the formation and evolution of the
83 Azores is supported by several authors. The Azores plateau and nearby topographic anomalies
84 were first attributed to a mantle hotspot (Schilling, 1975) and its interaction with the Mid-
85 Atlantic Ridge. Gente *et al.* (2003) proposed the Great Meteor Seamounts as the original
86 location of the hotspot, 85 Ma ago, before migrating northward, and showed that the
87 development of the plateau occurred 20 Ma to 7 Ma BP, before being tectonically rifted. Body
88 wave tomography by Yang *et al.* (2006) mapped a low-velocity anomaly beneath the Azores
89 Plateau at approximately 38.5°N, 28.5°W, extending northeastward and downward to connect to
90 a plume-like column from ~250 km to at least 400 km depth, centered northeast of Terceira
91 Island. Cannat *et al.* (1999) emphasized the influence of the Azores hotspot on the buildup of
92 the morphology of the MAR close to the Azores, and interpreted the Faial Ridge and a number
93 of other volcanic ridges as the result of a northward mantle flux along the MAR. Vogt and Jung
94 (2003) and Beier (2006) interpreted the existence of volcanic ridges south of Terceira Rift as
95 volcanic growth along a fossil rift and attributed the northward progression of the active rifting
96 area as a consequence of differential motion of the hotspot with respect to the overlying
97 lithosphere.

98 Space domain studies of the geoid to topography ratio in the Azores, found that the Plateau may
99 be supported solely by thickened crust (Grevemeyer, 1999). The same conclusion was achieved
100 by the work of Luis and Neves (2006) that investigated the gravity and geoid to bathymetry
101 relationship in the spectral domain. They found that the anomalously shallow depths of the
102 Azores Plateau were not due to dynamic forces sustained by mantle upwelling. The alternative
103 explanation was that the plateau is supported by thickened crust, which mainly results from
104 large volumes of accreted extrusives.

105 **Morpho-structure of the Azores Plateau**

106 The first regional bathymetric compilation of the Azores region was made by Krause and
107 Watkins (1970) and this supports their plate tectonic interpretation of the Azores. Searle (1980)
108 added new vertical soundings to the surveys made by Laughton *et al.* (1975) in the area and a
109 large amount of GLORIA side-scan sonar data along the plateau. The first systematic swath
110 bathymetry survey of the Mid-Atlantic ridge, close to the Azores, was made by Needham *et al.*
111 (1991) up to a distance that corresponds roughly to MAR magnetic chron c3 (~5 Ma). Lourenço
112 *et al.* (1998) extended Needham *et al.* (1991) compilation to the Azores, combining vertical
113 soundings with swath bathymetry data. Surveys over the MAR southern section were extended
114 to a wider off-axis region (Cannat *et al.*, 1999) and to the Azores triple junction, northwards off
115 the Azores plateau (Goslin *et al.*, 1999). Gente *et al.* (2003) collated the data originally prepared
116 by Thibaud *et al.* (1998), into a 1 km grid, covering the area 32-49N and 22-43W. While not
117 homogeneous, and not as detailed close to the islands as the compilation by Lourenço *et al.*
118 (1998), it gave the first regional overview covering the entire Azores Plateau and the MAR.
119 Nevertheless two key areas were not well described in these compilations: the complex tectonic
120 area close to the triple junction and the succession of basins that characterizes the Terceira rift.

121 Luis *et al.* (2006) made a high quality swath bathymetry survey of the Azores Triple Junction
122 area, using the French vessel Le Suroît and a Kongsberg EM300 multibeam echosounder. This
123 system images a variable swath width with 135 beams and a 32-kHz pulse, reaching a
124 maximum swath of 140°. For the average depth of the surveyed area (less than 2000 m below
125 sea level), this configuration enabled a typical 4-km across-track swath to be covered and a
126 50x50m grid to be produced. Two areas were surveyed in 2006 and 2007 (see Fig. 2 for
127 locations) covering essentially an area between the plateau and the MAR. The area is bounded
128 to the north by the North Azores Fracture Zone (~39°30'N) previously identified by Searle
129 (1980) as the probable structure connecting Terceira Rift with the MAR. The southerly limit is
130 bounded by latitude 38°30'N, which was identified by Luis *et al.* (1994), based on magnetic
131 data, as the probable location of the Azores triple junction.

132 Terceira rift was surveyed in the course of the United Nations Convention on the Law of the Sea
133 (UNCLOS) Surveys, by the Portuguese NRP Gago Coutinho, equipped with Kongsberg EM120
134 and EM170 systems for deep and intermediate to shallow water depths respectively (Madureira
135 *et al.*, 2011). A 50x50m grid was produced covering an area of approximately 15400 km².
136 Dedicated surveys were processed following IHO order 3 standard for hydrographic surveys
137 (IHO, 1987). Transits of opportunity were corrected using climatological profiles and added to
138 the systematic surveys as additional coverage to fill in the data gaps.

139 In Fig. 2 we present a 250 m bathymetric grid collating all new data. Most of the large scale
140 characteristics known from previous studies (Needham *et al.*, 1990; Lourenço *et al.*, 1998;
141 Gente *et al.*, 2003) are here expressed: the topographic anomaly which affects the MAR
142 between 35°N and 44°N generates a triangular plateau approximately bounded by the 2000 m
143 isobaths, and split by the MAR between a small western plateau and a large eastern plateau. Its
144 southern limit is marked by a major fracture zone (East Azores Fracture Zone) and the sequence
145 of steep basins that reach 2500-3000 m deep and volcanic highs which form the Terceira rift
146 developed close to its northeastern flank. Southwest of the Terceira rift one can observe the
147 volcanic highs of S. Jorge, Pico and Faial islands, and the Princess Alice Basin and bank, all
148 sub-parallel to Terceira rift. A number of small elongated volcanic ridges punctuate the plateau
149 with a similar strike.

150 The bathymetric data presented in Fig. 2 provides an accurate description of abyssal hill
151 topography associated with rifting at the MAR and with rifting at Terceira axis. It is the result of
152 high-angle faulting in slow spreading ridges and accounts for 10-20% of plate separation (cf.
153 Searle and Escartin (2004) and references herein). The geometry of abyssal hills and their
154 distribution can be used to identify the different tectonic patterns present in the Azores Plateau,
155 and in some situations they can be interpreted as chrons similar to what is done with marine
156 magnetic anomalies.

157 The small thickness of the sediment cover (Gente *et al.*, 2003) and the near absence of volcanic
158 features, allows for a clear bathymetric expression of the Azores related tectonic pattern. The
159 area between the MAR and 28°40'W, which corresponds to the northwestern tip of the Western
160 Graciosa Basin, is dominated by MAR-related abyssal hill topography, with a strike similar to
161 the MAR: ~N12E. At this stage there is no evidence of a single structure relaying Terceira Rift
162 to the MAR.

163 East of 28°40'W, the Terceira Rift becomes the main structural entity that disrupts the Azores
164 Plateau. MAR-related abyssal hill topography is dissected by ~N140° high angle faults
165 associated with the extension of the Western Graciosa Basin. The fault strike rotates
166 continuously up to N110°-120° which corresponds to the direction of the flanks of the eastern
167 and western Graciosa basins, and of the islands located south of the Terceira Rift. Basin flanks
168 are uplifted as a consequence of the equilibrium between isostasy and rifting extension (Lin and
169 Parmentier, 1990). The North Hirondele basin flanks strike ~N130° while the South Hirondele
170 basin flanks strike ~N140°. The Povoação basin, the easternmost basin of Terceira rift strikes
171 ~N160°, almost orthogonal to the small circle of the Eurasia-Nubia present-day infinitesimal
172 rotation pole, as given by MORVEL model (DeMets *et al.* 2010). Between rift basins, three
173 large volcanic structures form the islands of Graciosa, Terceira, and S. Miguel, and shallow
174 banks such as D. João de Castro (between North and South Hirondele basins) and the Formigas

175 islets (east of the Povoação Basin). S. Miguel is differently located: the western part of the
176 island is aligned with the South Hironnelle Basin, while the eastern part of the island is aligned
177 with Povoação Basin. Taken together this succession of deep basin and topographical highs
178 defines the crude segmentation pattern of the Terceira rift which presents a typical wavelength
179 of 60 to 100 km between major volcanoes or main basin depocenters along strike.

180 Close to the southwest corner of the eastern Azores Plateau, there is another steep basin that
181 shares strong similarities with the Terceira Rift: the Princess Alice Basin. It offsets the Faial
182 Ridge, approximately 20 km, at its eastern end. It is bounded by steep N130° flanks and
183 approximately 1200m deep, when compared to the neighboring plateau.

184 **Magnetic Analysis**

185 Magnetic surveys and magnetic chronology

186 The first magnetic compilation of the Azores was prepared by Krause and Watkins (1970) and
187 used to support the existence of a secondary spreading centre. In the late eighties, a large
188 aeromagnetic survey covered the entire plateau. First results were presented by Miranda *et al.*
189 (1991) and the whole survey was processed and interpreted by Luis (1996) and Luis *et al.*
190 (1994). The survey was dense and homogeneous but suffered from the limitations of the inertial
191 positioning system used for the aeromagnetic survey. Luis and Miranda (2008) presented a new
192 grid 2' magnetic compilation of the North Atlantic between 35°N and 47°N, where the most
193 accurate (post-GPS) marine data were collated. These magnetic data were merged with
194 complementary magnetic profiles acquired in the course of the UNCLOS Surveys. The final
195 grid was continuously reduced to the pole following the technique described by Luis and
196 Miranda (2008). In Fig. 3 we present the magnetic anomaly grid reduced to the pole, with a
197 resolution of 30'' covering the area 36N-41N and 32W-24W. Magnetic chrons were picked
198 from the correlation between measurements and synthetic profiles along small circles of the
199 presumed relative finite rotation poles. The following chrons were identified: C2, C2a, C3, C3a,
200 C4, C4a, C4, C5, C6, C6c, C13, C18, C20, C21, C22, C23, C24, C25, C26 and C28.

201 The interpretation of the magnetic anomaly map is straightforward: between the Mid-Atlantic
202 Ridge and chrons C4a-C5 magnetic striping is normally developed. Magnetic anomalies are
203 continuous between the North Azores Fracture Zone and Princess Alice Basin and spreading is
204 mostly regular since chron C4a. South of 38°30'N there is evidence of a jump to the west of the
205 ridge axis in the segment 38°N-38°30'N at chron C3, associated with the northward progression
206 of the triple junction (Luis *et al.*, 1994).

207 Between chrons C5 and C22 MAR magnetic chrons are disrupted by a sequence of linear
208 magnetic anomalies sub-parallel to Terceira rift, but with no systematic magnetic stripping.

209 These anomalies have high amplitudes (~1000 nT) and so probably correspond to the youngest
210 magnetic ages of Brunhes and Matuyama. The strike of magnetic anomalies changes
211 continuously from east to west, between Gloria Fault and the western Graciosa Basin: South
212 Hirondele Basin ~N125°, North Hirondele Basin ~N135°, East Graciosa Basin ~N125°, West
213 Graciosa Basin ~N140°. The only place where magnetic stripping seems better developed
214 corresponds to the East Graciosa Basin close to Terceira Island, where a succession of reduced-
215 to-the-pole positive and negative anomalies can be found. The two positive anomalies that
216 follow the southern and northern flanks of the rift basin and can be interpreted as magnetic
217 chrons (C2?). This is the oldest magnetic signature associated with the development of the
218 Terceira Rift.

219 South of the Terceira Rift linear magnetic anomalies match the volcanic highs of S. Jorge, Pico
220 and Faial islands with strikes ~N110°-N120°. They are not a topographic effect as shown by
221 inversion (Luis, 1996). In several places these anomalies continue on-shore where their ages
222 have been determined by radiometric methods. This is the case in the islands of Pico, Faial and
223 S. Jorge, where K/Ar ages given by Féraud *et al.* (1980) are compatible with the Brunhes-
224 Matuyama transition interpreted on magnetic anomalies (Miranda *et al.*, 1991). Radiometric
225 ages determined for Faial (Hildenbrand *et al.*, 2011), S. Miguel (Johnson *et al.*, 1998) and S.
226 Jorge (Silva *et al.*, 2012) reinforce the interpretation that the Matuyama-Brunhes transition is
227 found almost everywhere in the Azores islands and that Azorean magnetic anomalies are
228 younger than ~2 Ma. The exception is Santa Maria Island, believed to have formed between 5.2
229 Ma and 4.6 Ma (Féraud *et al.*, 1981), and the only island of the Azores archipelago where
230 fossiliferous sediments of Zanclean age (5.3Ma-3.6Ma) were found (Abdel-Monem *et al.* 1974;
231 Janssen *et al.*, 2008).

232 Princess Alice basin also disrupts the MAR-related magnetic pattern and its western limit
233 follows chron c2a (~3 Ma) at 38°30'N. It is very difficult to identify magnetic lineations on the
234 basin floor.

235 Islands Aeromagnetic Surveys

236 In 1985 two aeromagnetic surveys were made in the Azores islands covering Faial and S.
237 Miguel, respectively. The results of the first were published by Miranda *et al.* (1991) and
238 showed the existence in SW Faial of the Brunhes-Matuyama onshore. A recent work
239 (Hildenbrand *et al.*, submitted) that included extensive analysis of radiometric ages of volcanic
240 outcrops in the main units of Faial, and complementary paleomagnetic research, confirmed the
241 existence of at least two main phases in the construction of Faial Island. The older phase is
242 associated with widespread volcanism older than 800 ka, and consequently magnetically
243 reversed. The younger phase is directly magnetized, and extends up to the present-day.

244 The aeromagnetic survey of S. Miguel was made in 1985 at an average altitude of 1200 m
245 (western panel) and 1500 m (eastern panel). The magnetic anomalies, after analytical
246 continuation to a common level of 1500 m, and subtraction of the International Geomagnetic
247 Reference Field for the corresponding epoch, are plotted in Fig. 4A. There is a general
248 agreement between the plateau survey and the more accurate island aeromagnetic survey, and
249 one can observe that most of the island displays high amplitude positive magnetic anomalies,
250 coherent with the existence of highly magnetized and young eruptive rocks covering the flanks
251 of most of the volcanic systems. Two of the calderas (Sete Cidades and Furnas) match relative
252 magnetic lows and Fogo Volcano flank corresponds to a negative magnetic anomaly. A similar
253 situation can be found in the Nordeste area, with a WNW-ESE magnetic negative trend.

254 Using the total magnetic anomaly field and the bathymetry of S. Miguel, an equivalent
255 horizontal distribution of the magnetization is computed using some assumptions on the
256 thickness and vertical homogeneity of the magnetic layer. This inversion procedure, developed
257 by Parker and Huestis (1974), eliminates the dependence of the magnetic anomaly field from
258 the effect of the topography and the declination and inclination of the main field and the
259 magnetization. We considered a constant thickness source layer of 0.5 km, whose upper surface
260 follows the bathymetry, and a geocentric dipole magnetization with inclination 54° and
261 declination -13° . The equivalent magnetization draped upon the topography is shown in Fig. 4B.
262 It emphasizes the magnetization decrease associated with the calderas of Sete Cidades and
263 Furnas, the large negative magnetization associated with Fogo Volcano and linear negative
264 magnetization associated with the Nordeste Volcanic Complex.

265 We interpret the magnetization decrease as the result of hydrothermal alteration. A similar
266 interpretation was given by Blanco *et al.* (1997) for the decrease in the magnetic field anomaly
267 of Furnas, detected by a land magnetic survey. The intense fumarolic and soil degassing activity
268 observed at Furnas caldera (Ferreira *et al.*, 2013-this issue; Viveiros *et al.* 2013-this issue)
269 suggests the presence of a mature hydrothermal system beneath this central volcano and
270 supports such interpretation. On Sete Cidades Volcano present-day degassing processes are not
271 so evident due to the presence of a large superficial aquifer system but the intracaldera volcanic
272 activity in the last 5000 years shows the largest eruptive frequency in the Azores suggesting a
273 very active plumbing system (Queiroz *et al.*, 2008; Queiroz *et al.*, 2013-this issue). The negative
274 magnetization of Fogo Volcano and Nordeste Volcanic Complex are probably the result of
275 reversed magnetization, acquired before the Brunhes-Matuyama transition, as is found
276 elsewhere in the Azores. In the case of Fogo Volcano, it can be the consequence of a thick layer
277 of reversely magnetized rocks, partially overlain by more recent volcanic products (Wallenstein
278 *et al.*, 2013-this issue), which (vertically integrated) effect can be seen in Fig. 4B. In the case of

279 Nordeste Volcanic Complex this interpretation matches the radiometric ages determined by
280 Johnson *et al.* (1998) in surface outcrops.

281 **Present-day surface displacement field**

282 Inter-island displacement field

283 The first GPS network campaign was made in the Azores in 1988 that included only one station
284 per island. Observations have been carried regularly since then (see Fernandes *et al.* 2006 and
285 references herein).

286 A number of studies have been based in campaign GPS observations. Fernandes *et al.* (2004,
287 2006) analyzed a set of 9 stations, corresponding to the period 1993-2000 and 1993-2001. They
288 concluded that the absolute motion for all stations was intermediate between pure Eurasia and
289 pure Nubia behavior where the end members appeared to be Graciosa and Santa Maria
290 islands, respectively. All other sites were located along the presently active inter-plate
291 deformation zone. This was also supported by the averaged motion per island using the
292 solutions from the dense networks (Fernandes *et al.*, 2006; Miranda *et al.*, 2012).

293 More recently, the number of permanent GNSS stations in Azores has also improved
294 significantly. This work, we focuses only in such type of stations, which although still with a
295 distribution less dense than the existing networks of episodic stations (particularly on the islands
296 of Terceira, S. Jorge and Faial), are already able to provide reliable estimates concerning the
297 dynamics of the region. In fact, such stations do not suffer from reoccupation errors (for most of
298 the Azorean permanent stations, the antenna and receiver have been the same since the initial
299 installation) and most of the effects due to seasonal signals have been already mitigated after
300 2.5yr (Blewitt & Lavallé, 2002). Table I details the most significant properties of the network of
301 permanent stations in Azores available to this research group. In addition to the stations with
302 data-span smaller than 2.5yr (VFDC and PTRP), TERC has also been excluded from the
303 analysis. This station has many data gaps and a significant number of outliers in a relative short
304 time-series (2.9yr).

305 Fig. 5 shows the estimated velocity field with respect to stable Nubia using as reference the
306 SEGAL08 plate tectonic model (Fernandes *et al.*, 2011). This model includes estimations of the
307 angular velocity of Nubia and Eurasia derived from GPS velocities computed with respect to
308 ITRF08, the latest realization of the International Terrestrial Reference System. For comparison,
309 we also plot the predicted relative motion for points in the stable Eurasia using SEGAL08,
310 GEODVEL (Argus *et al.*, 2010) and MORVEL (DeMets *et al.*, 2010). GEODVEL is also a
311 space-geodetic based angular velocity model whereas MORVEL is the most recent
312 geophysical/geological angular velocity model averaging the motions of the plate tectonics in

313 the last 3Ma. As it is observed, the predicted motions given by all these models are similar with
314 SEGAL08 and GEODVEL (N82°-85°E) predicting a more oblique opening of the Terceira Rift
315 – almost West-East, whereas MORVEL (N69°-77°E) shows a more pure opening regime of the
316 Nubia-Eurasia plate boundary between S. Miguel and Graciosa islands.

317 The Graciosa station (AZGR) presents an abnormal residual suggesting this island is presently
318 locked to Eurasia. Even if the large uncertainty of the derived solution does not allow a robust
319 conclusion to be reached, this estimation is coherent with other solutions computed in the past
320 based on episodic data (Fernandes *et al.*, 2004, 2006), which also showed Eurasia motion for
321 this island.

322 The results for the other stations in Faial, Terceira and Pico (FAIM, TOMA and PIED,
323 respectively) also confirms the previous computed velocity fields of these islands presented by
324 Fernandes *et al.* (2006) and Miranda *et al.* (2012) based on the combination of permanent and
325 episodic data (for Faial and Terceira) or episodic data only (Pico). Clearly, Faial and Pico
326 present a motion that is practically Nubia whereas the motion for Terceira is intermediate
327 between Nubia and Eurasia.

328 The vertical component of the surface displacement field is less constrained than the horizontal
329 components. There is evidence for subsidence of Terceira Island close to 0.8 mm/yr . PDEL
330 station in S. Miguel Island shows also subsidence rate of -0.66 +/- 0.21 mm/yr in the last 10
331 years of continuous observations (Miranda *et al.*, 2012). The station showing the largest
332 subsidence is also located in S. Miguel – FRNS (-3.21 +/- 1.51 mm/yr). In fact, the only station
333 showing uplift is AZGR (1.46 +/- 2.73 mm/yr) but the associated uncertainty is too large to
334 derive any sound conclusion.

335 Intra-island deformation field

336 Intra-island deformation was studied in S. Miguel Island (Sigmundsson *et al.*, 1995; Jónsson *et al.*
337 *et al.*, 1999; Trota *et al.*, 2006; Trota, 2008), Terceira (Navarro *et al.*, 2003, Trota, 2008, Miranda
338 *et al.*, 2012), Faial (Catalão *et al.*, 2006), and Pico (Catalão *et al.*, 2010).

339 Jónsson *et al.* (1997) analysed GPS campaign data from S. Miguel Island and found significant
340 extension within the island. They hypothesize that a 10-15 km wide zone, in the Furnas region
341 was accommodating about 75% of the divergence between Eurasia and Nubia plates. Trota *et al.*
342 (2006) extended this study, analysing data from five GPS campaigns between 1993 and 2002 in
343 S. Miguel. They concluded that a somewhat larger zone of the island, ~15-25 km wide,
344 accommodates most of the deformation WNW–ESE extension at a rate close to 5 mm/yr ,
345 which has the same order of magnitude as the Eurasia-Nubia relative motion according to all
346 geodetic models.

347 Our results based on the two analysed permanent GPS stations in S. Miguel (PDEL and FRNS)
348 also confirm an extension of the island almost in the WNW-ESE direction (N100°E). However,
349 the estimated opening rate using the derived secular motions (2.2 mm/yr) is only about half of
350 the predicted opening rate given by the considered tectonic angular velocity models (cf. Figure
351 5). This opening rate is even reduced if only the common period of observations for both
352 stations is considered (September 2007 – November 2011) and only the rate of the change of the
353 baseline between both stations is computed (which is independent of any errors associated with
354 the reference frame). In this case, the averaged opening rate is reduced to 1.64 +/- 0.20 mm/yr.
355 This is only about 35% of the expected deformation.

356 Different results were obtained for Terceira Island, where no significant intra-island
357 deformation patterns were identified and the magnitudes of local deformation processes
358 measured by the GPS network are low when compared with the accuracy of the velocities.
359 According to Miranda *et al.* (2012) Terceira moved most closely with Eurasia during the period
360 1996-2010 and the two studies (Navarro *et al.*, 2003 and Miranda *et al.* 2012) point to small
361 lateral compression, with no evidence of extensional processes.

362 Intra-island deformation of Faial was studied by Catalão *et al.* (2006) using conventional
363 geodetic measurements made between 1937 and 1997 and GPS campaign data. They showed
364 that Faial intra-island deformation during most of the 20th century could be attributed to a single
365 event (Capelinhos eruption and subsequent earthquake), and not to tele-tectonic processes. A
366 similar conclusion was obtained for Pico, where surface displacements are related with mass
367 waste processes (Catalão *et al.*, 2010).

368 All existing studies (Catalão *et al.*, 2006, 2010) point to the importance of local volcanic
369 dynamics on the intra-island deformation field. Inflation leads to the development of elongated
370 volcanic ridges associated with all large volcanic units and deflation leads to contraction and
371 graben development (Miranda *et al.*, 2012).

372 **Discussion and Conclusions**

373 Vogt and Jung (2004) considered that while the distance between magmatic centers in the
374 Azores is similar to what is found in ultra-slow ridges, the rift valley is wider than the nearby
375 MAR, as a consequence of the smaller spreading rate (4 mm/yr compared with 23 mm/yr at a
376 latitude of 40°N), corresponding to a thicker axial lithosphere, geometrically similar to what is
377 found in extinct rifts (e.g. Guadalupe and Aegir). A similar situation is supposed to occur on the
378 amplitude of the topographic segmentation (2000-4000 m) which greatly exceeds that of
379 “normal” slow MAR segments and registers the more abundant magma supply that has created
380 the Azores Plateau. The data presented here do not contradict this view: ultra-slow spreading
381 along the Azores arm creates morphological highs and deep basins along the Terceira Rift, with

382 episodic volcanic events south of the main rifting area, focusing under large volcanoes instead
383 of occurring widespread along the Terceira rift zone.

384 The northward migration of the Azores triple junction, with the progressive attachment of
385 Eurasia lithosphere to Africa has been identified as an important process that shaped the Azores
386 plateau (Lourenço *et al.*, 1998). However, only two major rift features can be found in the
387 plateau. The first one is the Terceira Rift, which corresponds to the present-day location of
388 rifting processes in the Azores, as demonstrated by historical and instrumental seismic and
389 volcanic activity (Gaspar *et al.* 2013-this issue), magnetic, tectonic (Madeira *et al.*, 2013-this
390 issue) and space geodetic data. The second one is the Princess Alice basin and bank, which is
391 now located south of the triple junction area, and most probably corresponds to the previous
392 location of the rift. The shift between the two must have occurred previous to chron C2a.

393 Vogt and Jung (2003) and Yang *et al.* (2006) proposed the existence of a succession of NE
394 jumps of the oblique spreading axis between an initial location close to the Princess Alice Basin
395 and a final one close to the so-called North Azores Fracture Zone. These jumps are seen as
396 consistent with the existence of a relative stationary Azores mantle “hotspot” across which the
397 Gripp and Gordon (2002) model predicts a N247 direction velocity varying between 22 and 33
398 km/Ma (Vogt and Jung, 2003). As the distance between Princess Alice Basin and the Western
399 Graciosa Basin, along the direction N247 is close to 150 km, this will imply a time span of 4.5-
400 6.8 Ma. This is inconsistent with the data and interpretation presented here. So, the construction
401 of the volcanic ridges between the two rifts must be considered as having occurred almost
402 concomitantly with the Terceira rift formation as an expression of off-rift volcanism and cannot
403 be interpreted as an intermediate stage materializing progressive migration of the Azores Triple
404 Junction.

405 The relative motion between Eurasia and Nubia along the Azores is mostly concentrated on
406 Terceira Rift as demonstrated by magnetic and geodetic data. While Graciosa and Terceira are
407 progressively attached to Eurasia, and subsiding as they move away from the rift axis, S. Miguel
408 is being strained by rifting with a significant amount of the inter-plate displacement field being
409 accommodated within the island. In S. Miguel subsidence is a more complex process and seems
410 to be particularly high in the central part of the island as it was first discussed by Muecke *et al.*
411 (1974) based on rock samples from a borehole drilled in the north flank of Fogo, where
412 subaerial rocks were found at about 800 m depth.

413

414 The identification of magnetic chrons between C2 and C5 shows that south of 38N Nubia
415 rotations apply, north of 39.5N Eurasia rotations apply, and between the two limits there is an

416 almost continuous change between the two (Miranda *et al.*, 2011). In this sense, there is no right
417 lateral strike slip fault along the North Azores Fracture Zone as hypothesized by Searle (1980)
418 or by Vogt and Jung (2003), but instead, the differential motion between Eurasia and Nubia,
419 west of Western Graciosa Basin is accommodated within a ~100 km wide subject to right lateral
420 transtension generating the complex tectonic pattern described above.

421 The angle between the Terceira basins flanks (or the strike of magnetic anomalies), and the
422 direction of the relative motion between Eurasia and Nubia as given by both geodetic and
423 “geological” kinematic plate models, was interpreted by Lourenço *et al.*, 1998 as “oblique
424 spreading”. However, both the morphological expression of the rift flanks and the trend of
425 magnetic anomalies point to the conclusion that rifting is approximately orthogonal, close to the
426 rift. We hypothesize that mantle processes associated with the velocity anomalies mapped by
427 seismic tomography (Yang *et al.*, 2006) are constraining the rift direction close to the actively
428 deforming zone. At the island scale, volcanic dynamics, in particularly inflation and deflation
429 control intra-island deformation and normal faulting.

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431 gave us the opportunity to acquire a significant part of the high quality swath bathymetry data
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433 the last 20 Ma between Kurchatov and Hayes, PTDC/MAR/108142/2008.

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603

604 **FIGURE CAPTIONS**

605 **Figure 1:** The Azores plateau defined by the 2000 m isobath. MAR: Mid-Atlantic Ridge; WGB:
606 Western Graciosa Basin; EGB: Eastern Graciosa Basin; NHB: North Hirondele Basin; SHB:
607 South Hirondele Basin; PB: Povoação Basin; PAB: Princess Alice Basin. Vectors at the flanks
608 of Terceira Rift basins indicate the relative motion between Eurasia and Nubia from geodetic
609 models

610 **Figure 2:** 250 m bathymetric grid. New swath bathymetry surveys are indicated by thin white
611 polygons.

612 **Figure 3:** Magnetic compilation and isochron identification. Chrons identified close to the Mid-
613 Atlantic Ridge are c2, c2a, c3, c3a, c4 and c4a. Other chrons are labeled. Black dots indicate
614 positive identification and thin solid black lines indicate the probable geometry of the isochron.

615 **Figure 4:** Velocities relative to Nubia, as given by the permanent GNSS stations. In spite of the
616 still large error associated with the determination of Graciosa velocity, most of the extension
617 takes place within Terceira rift and São Miguel Island shows significant intra-island extension.
618 Within the Eurasia plate, away from the boundary we plot the velocities given by models
619 MORVEL (violet), yellow (GEODVEL) and SEGAL (green). See text for details.

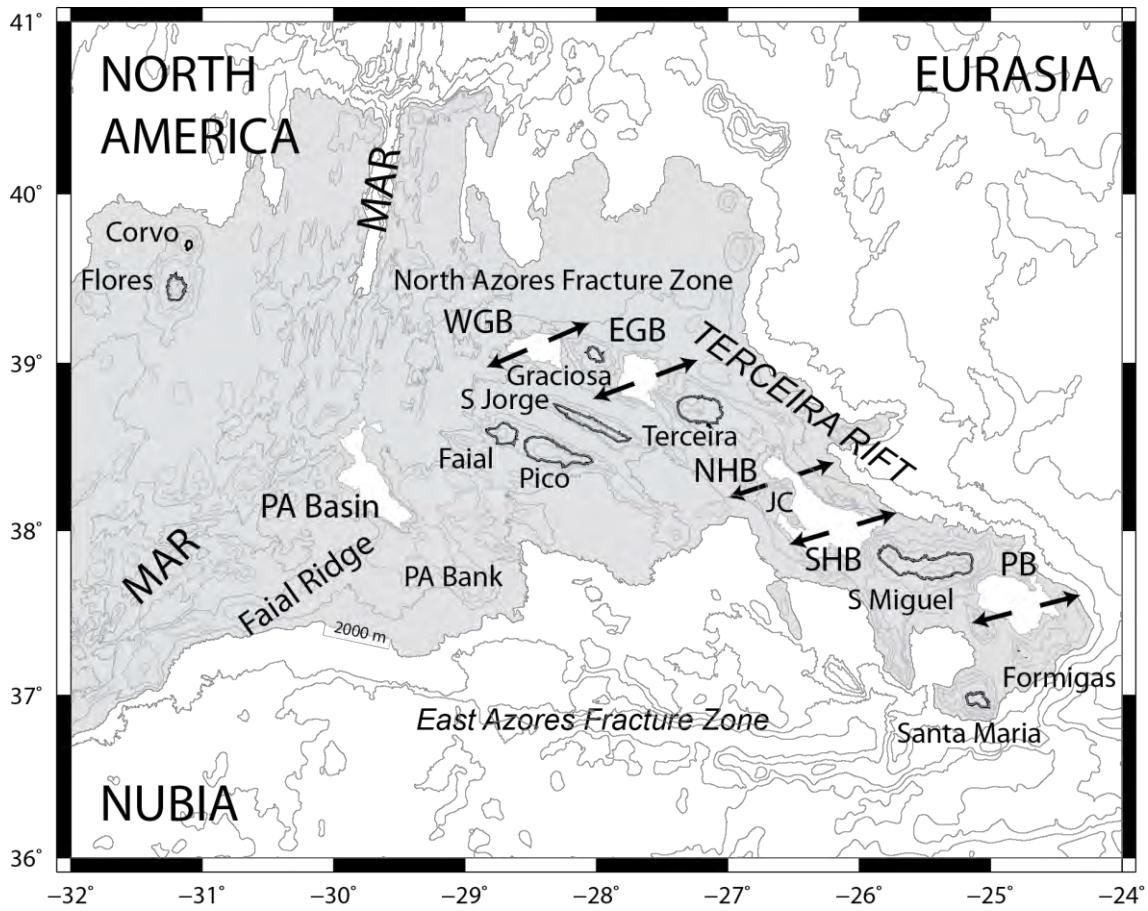
620 **Figure 5:** A: aeromagnetic Survey of São Miguel Island; total field anomalies are plotted every
621 100 nT. B: drape of magnetization on a detailed topographic map of S. Miguel Island.
622 Magnetization values plotted every 1 Am⁻¹. The main calderas are labeled: Sete Cidades; AP:
623 Agua de Pau or Fogo; FU: Furnas; PV: Povoação.

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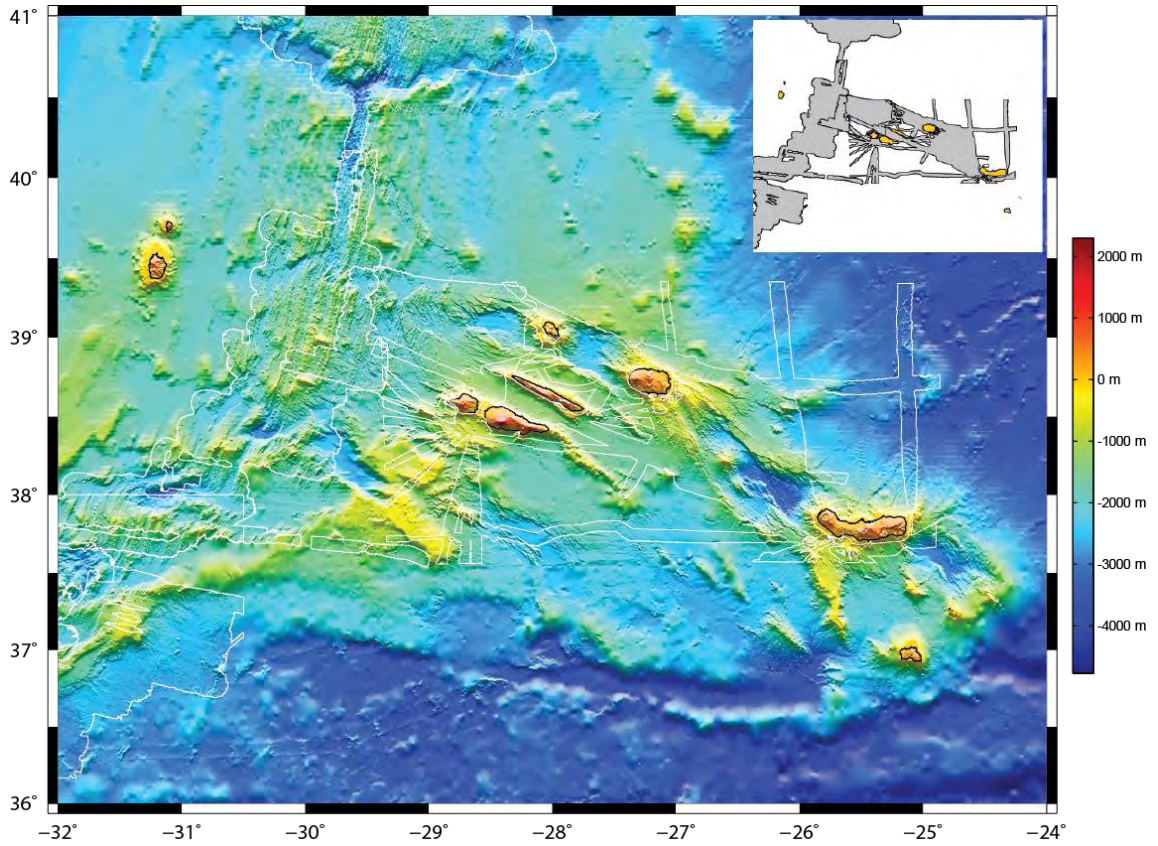
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629 Vectors at the flanks of Terceira Rift basins indicate the relative motion between Eurasia
630 Nubia from geodetic models.

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633 Fig.2. 250 m bathymetric grid. New swath bathymetry surveys are indicated by thin white
634 polygons.



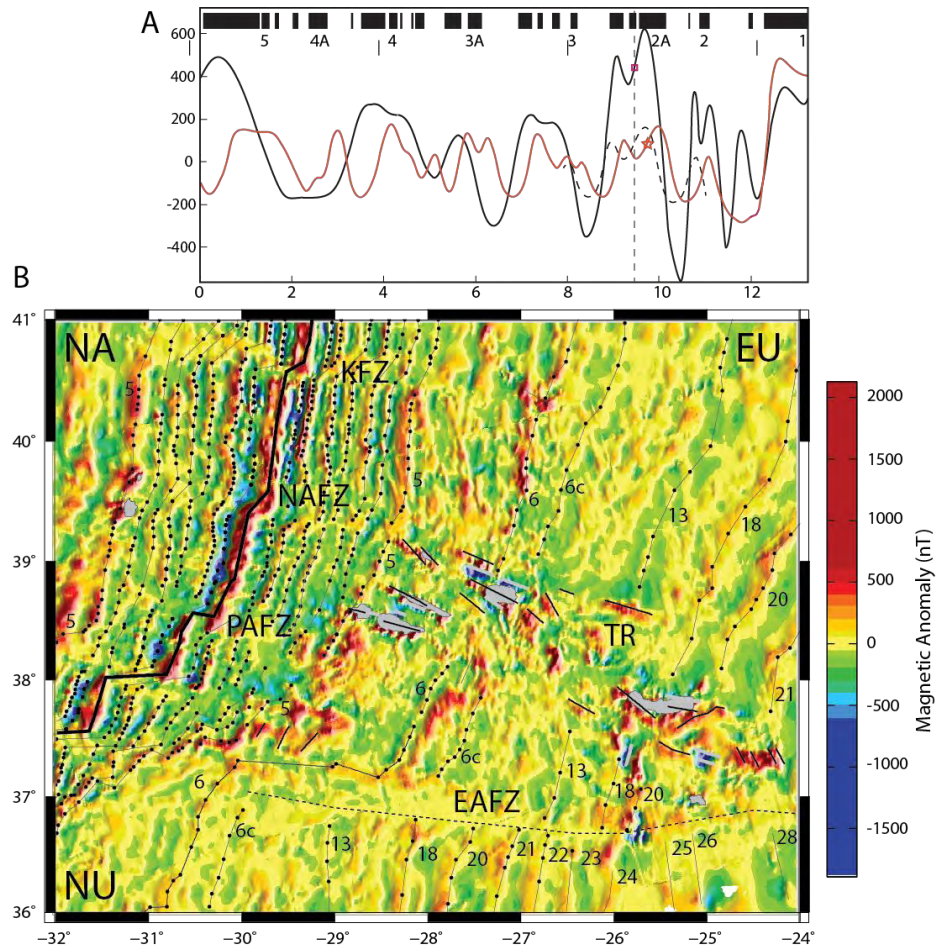
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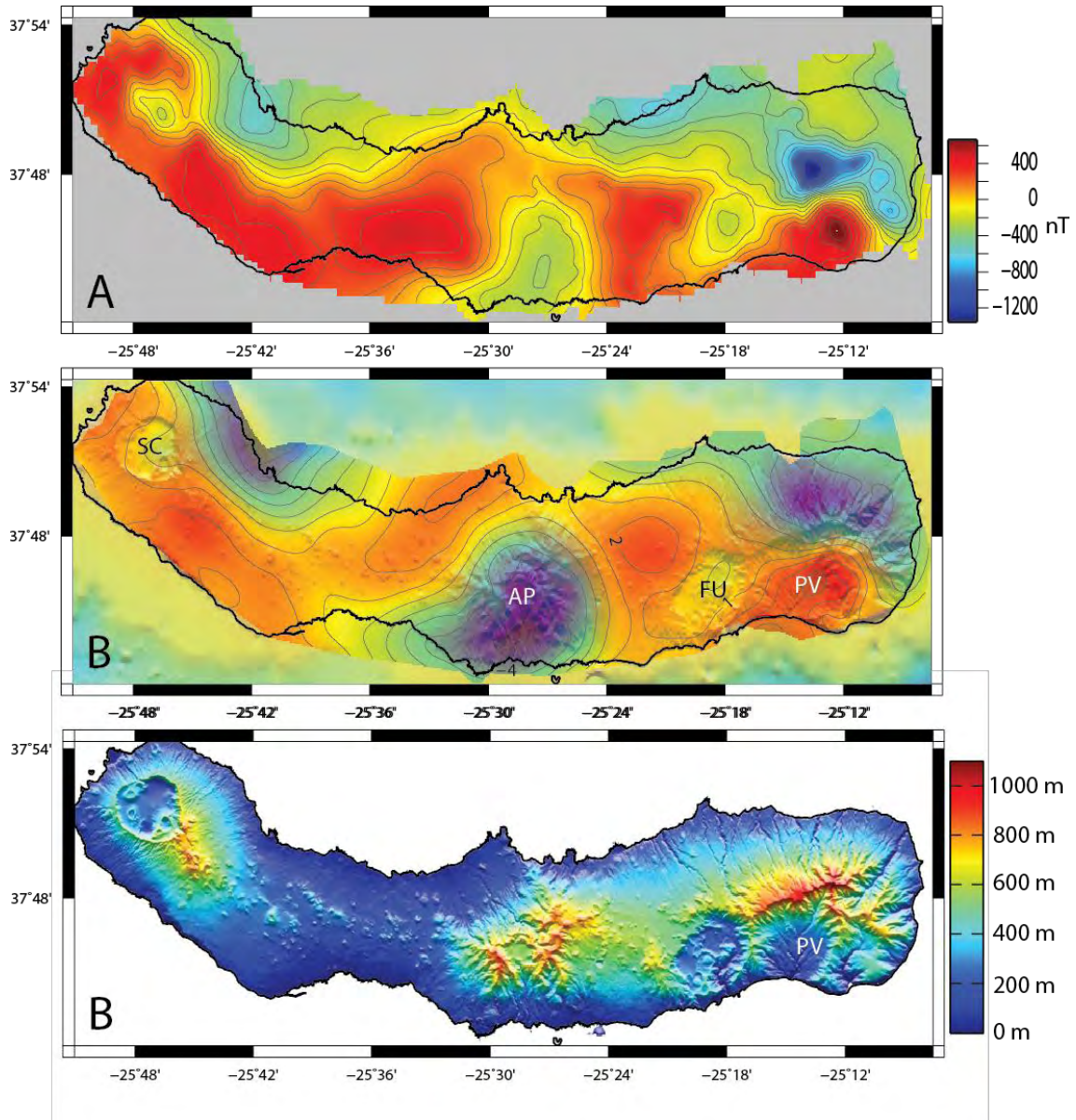
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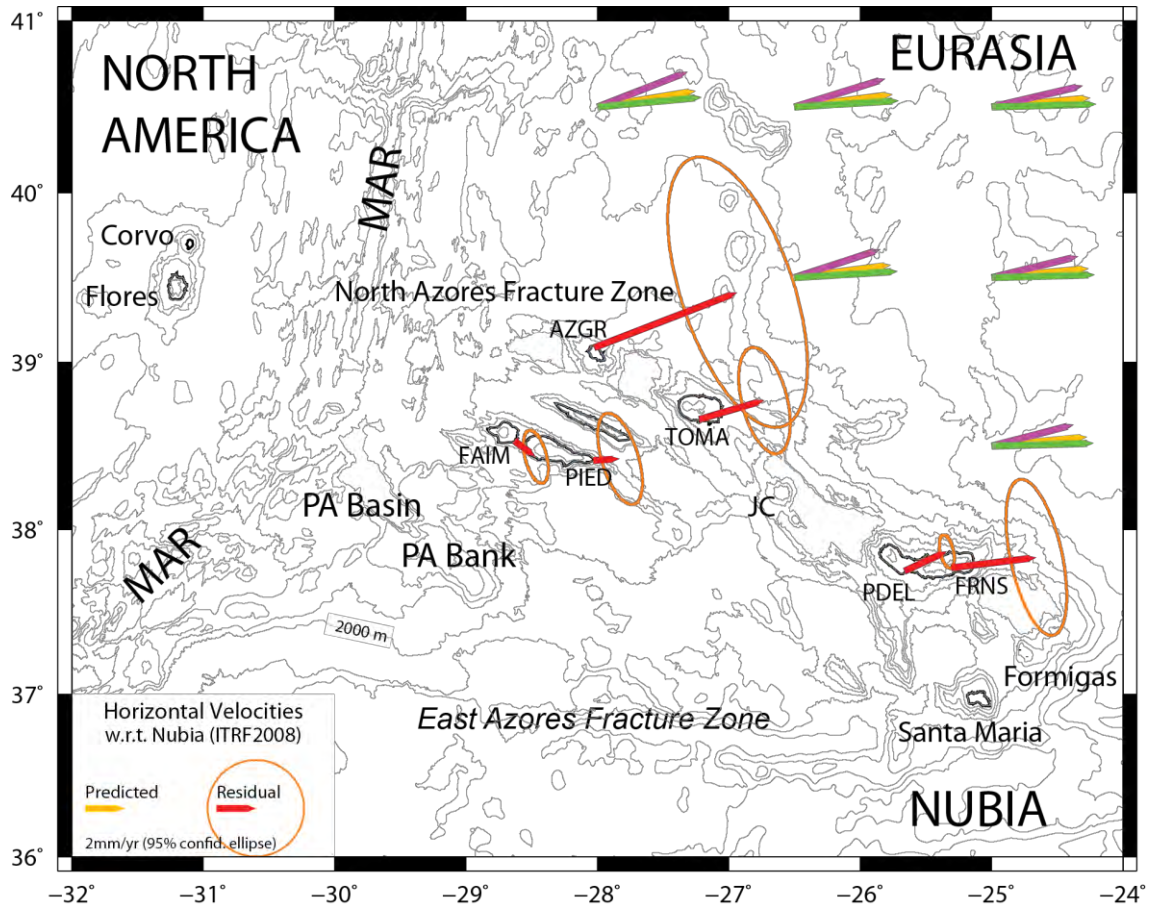
644 Fig.4. Top: aeromagnetic Survey of São Miguel Island; total field anomalies are plotted every
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 646 Magnetization values plotted every 1 Am^{-1} . Bottom: topographic map of S. Miguel Island. The
 647 main calderas are labeled: SC - Sete Cidades; FG – Fogo Volcano; ; FU - Furnas; PV -
 648 Povoação.



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651 Fig.5. Velocities relative to Nubia, as given by the permanent GNSS stations. In spite of the still
 652 large error associated with the determination of Graciosa velocity, most of the extension takes
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 656 João de Castro bank.



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660 Table 1. Permanent GNSS stations in Azores (see Figure 5 for locations). The only present
661 inactive station is FAIM.

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SITE	Starting Date	Final Date	Years	Total Days	Island	Network
PDEL	01-02-2000	23-11-2011	11.8	4067	S. Miguel	IGS
FAIM	05-10-2000	19-07-2010	9.8	2259	Faial	IDL
TOMA	02-10-2001	26-04-2010	8.6	1479	Terceira	IDL
PIED	27-07-2006	23-11-2011	5.3	957	Pico	REPRAA
FRNS	15-07-2007	23-11-2011	4.2	855	S. Miguel	REPRAA
FLRS	14-07-2008	23-11-2011	3.4	852	Flores	IGS
TERC	14-12-2008	21-11-2011	2.9	689	Terceira	REPRAA
AZGR	14-01-2009	22-11-2011	2.9	859	Graciosa	REPRAA
VFDC	26-03-2010	22-11-2011	1.7	511	S. Miguel	REPRAA
PTRP	19-05-2010	22-11-2011	1.5	247	Pico	REPRAA

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